



What's in a name – the early days of JIVE

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JIV-ERIC Symposium, Diwngeloo

20 April 2015





1975-1980 context

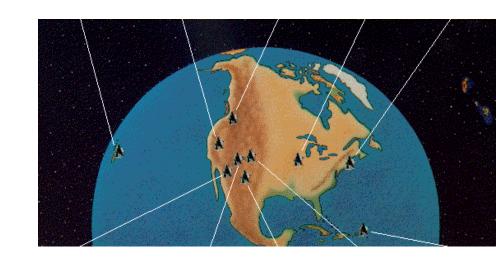


Genesis of the EVN see talk by Roy Booth

US Network established in 1976

NRAO proposal for an intercontinental array

See talk by Ken Kellermann







European discussions on (large) correlators

late 1970s Internal discussions in MPIfR, Jodrell Bank,

and NFRA

Nov 1979 MPIfR decision

- purchase 3-station Mk3 correlator from Haystack

Mar 1980 First meeting of Telescope Directors in Bonn

discussed satellite-linked VLBI and data processing

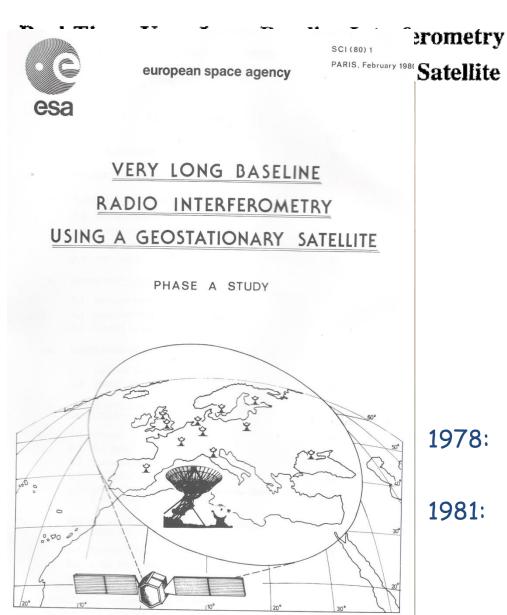
needs - 8-station real-time correlator



satellite-linked VLBI



1977



1978: ESA Feasibility Study of

satellite-linked VLBI

1981: ESA Phase A study of

satellite-linked VLBI using

L-SAT





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processing needs – 8 station real-time correlator

Mar 1980 – Exchange of letters between EVN (HvdL) and ESA on

May 1981 L-SAT opportunity





European discussions on (large) correlators (2)

early 1981 Proposal to utilise WSRT Digital Line backend design

for a VLBI correlator for satellite-linked VLBI

(Schilizzi, Miley, Goss)

Jun 1981 Director's Meeting in Leiden discussed EVN data

processor needs again

ESA required 2.5-3.5 MECU for modifications to L-SAT for VLBI + pay for ground stations at each telescope

→ Demise of satellite-linked VLBI

early 1982 QUASAT feasibility study begins

Jan 1983 Director's meeting in Garching

1) upgrade Mk3 processor at MPIfR to 12 stations

2) develop new generation (12 station) data

processor in Dwingeloo for the longer term future





The courtship of Brussels begins

Sep 1983 proposal by Giancarlo Setti et al to EC for

expansion of Mk3 correlator. Not successful

? 1983 Contact by Setti with Herbert Curien, President of

the ESF

Jul 1984 Directors Meeting in Vienna

European Consortium for VLBI formed with the aim to find funding for a large correlator centre

?1984 Setti contact with Prof Fasella, D-G of Research in

EC





The courtship continues for some time...

Feb 85

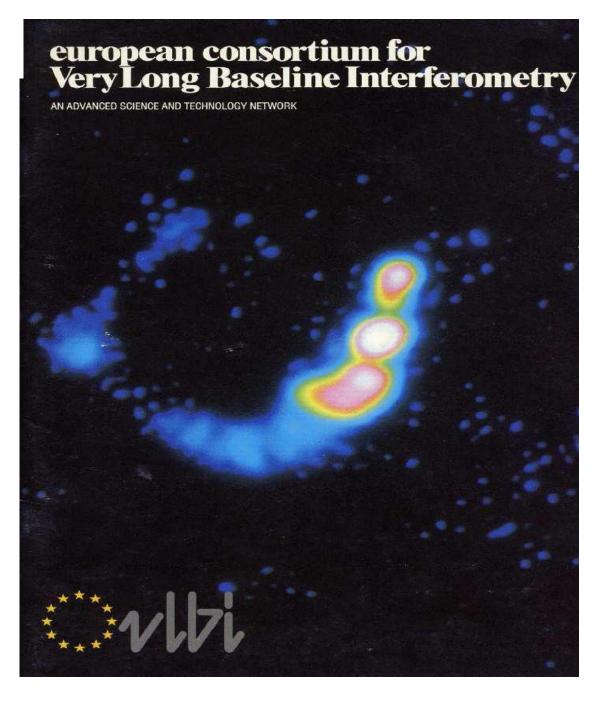
Decision by Consortium Board to seek funds from 'Europe' for a 12-station correlator to be located in Dwingeloo

MPIfR Directors decided to upgrade their 3-station correlator to 5-stations. (Operating a large correlator as a service facility for the European community not within MPG remit)

Nov 85

EVN brochure published











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Nov 85

EVN brochure published

May 86

visit to DGXII in Brussels by NL representatives, R. van Lieshout, Harry van der Laan, RTS





Serious proposal writing and "discussions"

Nov 86 Proposal for the EVN data processor (12)

stations) submitted to CEC and circulated to

European Science Ministries

Mar 87 – discussions with Prof Jules Horowitz, DGXII

Jan 88 consultant on Large Facilities Programme

Result: VLBI not recommended

Sep 87 EC Framework Programme approved

Jan + May 88 Consortium Board discusses "science" programme

with DGXII representatives

Jun 88 Proposal to EC for the first phase of a 20 station

data processor (total cost 17.8 M€)





Finally some traction

Dec 88 CODEST Committee report highlighted "the extremely high quality and scientific value of the project, but lack of funds in the science programme meant that only a feasibility study could be funded"

Consortium delegation discussed future strategy with Director of Science Stimulation Programme in DGXII (Herbert Allgeier)

Feb89 Nature uses VLBI as an example of a project that should be funded by the CEC





nature

NATURE VOL. 337 23 FEBRUARY 1989

675

What science for Europe's 1992?

An acid test of the European Community's intentions in support of science after 1992 is its willingness to respond to an imaginative proposal from radioastronomers.





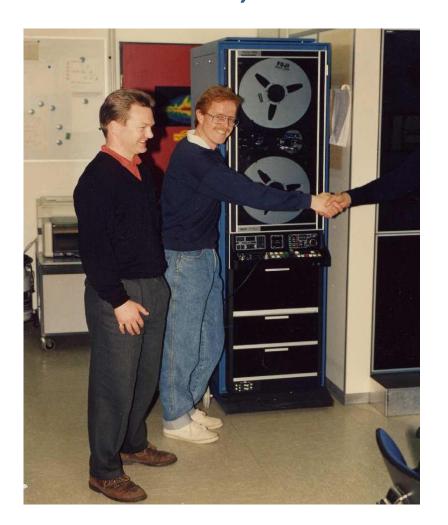
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Penny & Giles tape recorder delivered to Bonn, 1991







The final steps....

- Apr 90 Ministers agree to ESF Review NL and FR request ESF to take action
- Jul 90 ESF Review Panel on ground-based astronomy recommends funding by EC
- Oct 90 NL Ministry led pressure on Brussels for EC funding of pure research facilities like VLBI in the Framework programme
- May 91 NL Ministry reserves 12 Mfl (€5.5M) in the Budget for International Facilities (BIF)
- Sep 91 NL Ministry invites other European Science Ministries to a meeting to decide on VLBI funding





Funding at last!

19 Feb 1992

- 5.5 M€ from Ministry of Education and Science in NL
- 0.3 M€ from CNRS in France
- 0.55 M€ from the Swedish Wallenberg Foundation
- Engineering support from Jodrell Bank Observatory in the UK,
 CNR in Italy, and IGN in Spain
- ? 1992 FP3 HCM grant (1M€) to EVN for Access to Large Scale Facilities and \$300k for Fellowships.





What's in a name?

Why JIVE?

Easy to remember in an era of E-everything (ESO, ESA,EMBL,.. European VLBI Institute)

Not everyone approved!

African-American slang: deceptive, nonsensical, or glib talk

Jazz or swing music

Logo

Competition among the Dutch astronomy departments won by Olaf Kolkman from Groningen



9 June 1993

official opening of JIVE in Dwingeloo



Director of Science Policy in the NL Education and Science Ministry

Ambassador of Sweden Representatives of the Embassies of France, Italy, Spain and the UK Queen's Commissioner for Drente







mr. A.H.C. van Drooge notaris te Dwingeloo

AFSCHRIFT

ENER AKTE VAN

OPRICHTING

VOOR:

STICHTING "JOINT INSTITUTE for V.L.B.I. in EUROPE (J.I.V.E.)"



First Chair of JIVE Board – Roy Booth





The EVN Data processor at JIVE

1993 –1998

design, prototyping, and construction of EVN 16-station processor by international consortium

Part of the EVN Upgrade

Total cost

8.7 M€ including manpower

22 Oct 98

official opening of EVN Data Processor at JIVE





Correlator design

First meeting March(?) 1992

Albert Bos

Alan Whitney

Sergei Pogrebenko

Bryan Anderson

$$\nabla \cdot \mathbf{E} = \frac{\rho}{\varepsilon_0}$$

$$\nabla \cdot \mathbf{B} = 0$$

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$$

$$\nabla \times \mathbf{B} = \mu_0 \left(\mathbf{J} + \varepsilon_0 \frac{\partial \mathbf{E}}{\partial t} \right)$$



Responsibilities



Project Managers

Jean Casse (JIVE), Alan Whitney (Haystack)

Correlator design

JIVE – Albert Bos, Sergei Pogrebenko WSRT DXB– Albert Bos MkIV (Haystack, USNO, Bonn), SMA – Alan Whitney

Correlator chip

Will Aldrich, John Canaris

Station Unit

Bryan Anderson, Steve Parsley, Sergei Pogrebenko, Gino Tuccari, Stelio Montebugnoli

Playback drives

Steve Parsley, Jan Buiter

Software (JIVE)

Roger Noble (on-line), Huib van Langevelde (post-correlation)





Correlator chip story

- End-94, problem in finding a fabricator was solved by good luck
 - Alan Whitney met HP CEO, Lew Allen on a trip to Silicon Valley
 - HP foundry available, but only for 1 year
- Time for only one iteration
 - Test wafers, packaging and testing at Haystack in correlator board
- End-96, final production runs → 8000 chips, enough for the 4 correlators and spares





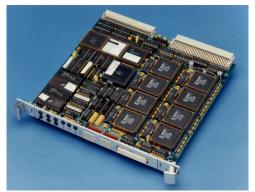
The correlator





Station Units













Playback drives



16 P&G playback drives





Correlator control software

Object oriented high level control software allowed relatively simple implementation of MkIII, VLBA and

MkIV modes







EVN support

Support Scientists

- on-site support to visiting astronomers
- planning, scheduling, analysis of data
- VLBI schools
- investigate
 calibration and
 polarisation of the
 EVN

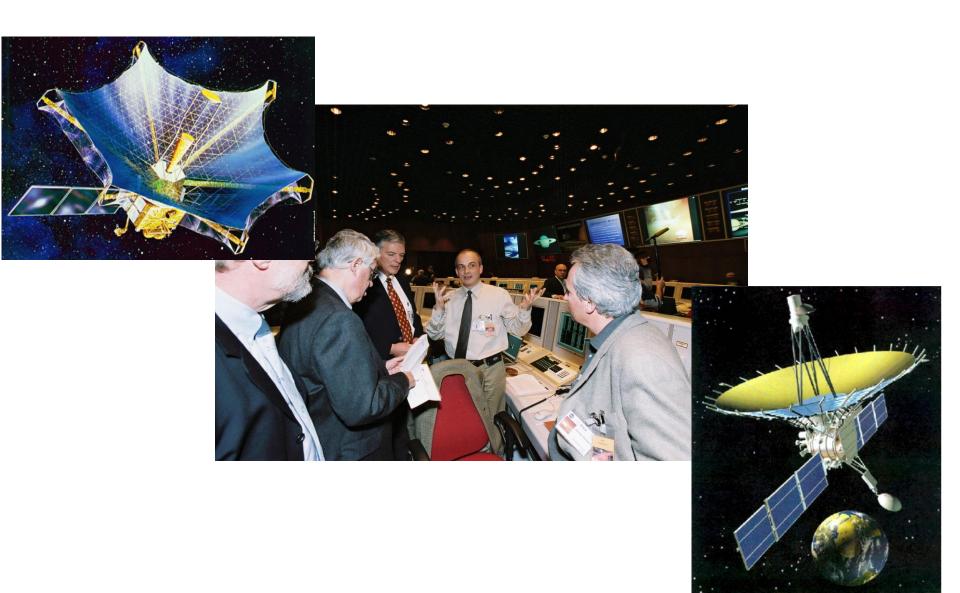
Human Capital & Mobility money being well used







Space VLBI

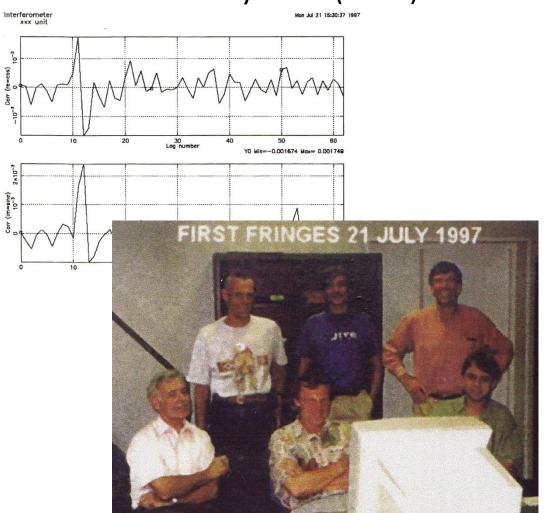




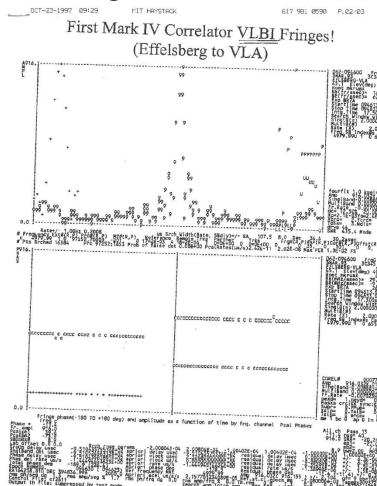


First fringes

JIVE on 21 July 1997 (MkIII)



Haystack first fully-tested VLBI fringes 20 Oct 1997





Official opening of the EVN Data processor at JIVE

22 October 1998

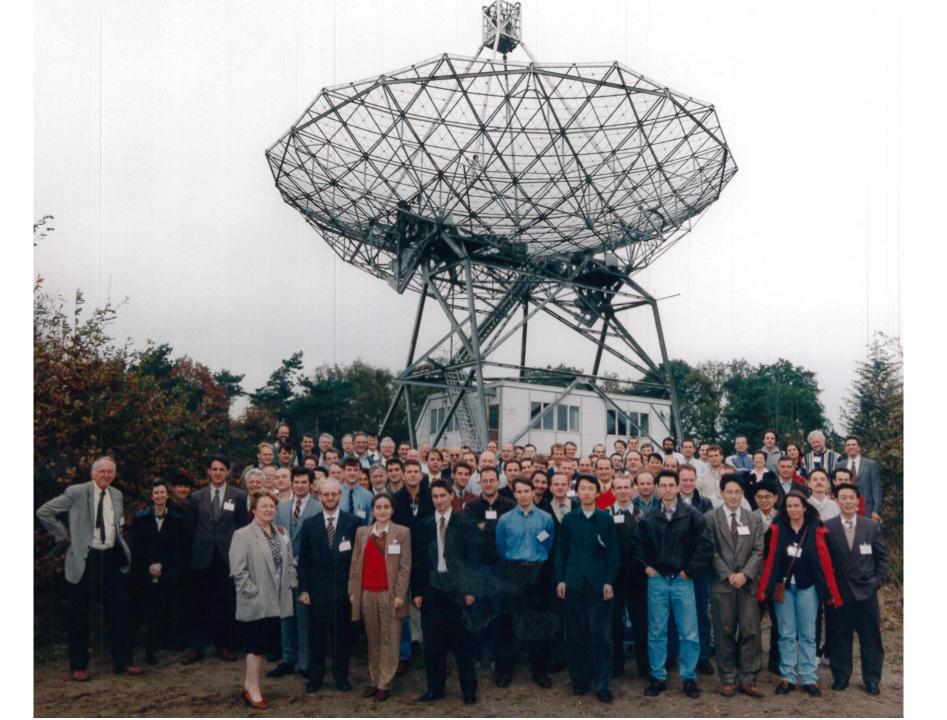


















Postscript

First EVN correlation in ~ July 1999

By 2000, essentially all EVN observations and ~50% of global observations correlated at JIVE





Correlator in operation







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2000 - first science publication

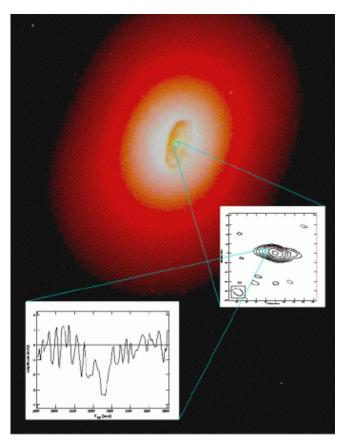




First science

van Langevelde et al 2000

"A thin HI circumnuclear disk in NGC4261"



The first scientific result of the SVN data processor at JIVE is locating a tonic Hydrogen close to the nucleus of NGC4851. The background is the Hubble telescope image showing a neutral accretion dask in the inner part of this active galaxy. Insets are the VLBI image of the nucleus at 21 cm (right) and the NLabsorption apportrum, observed slightly offset from the nucleus [left].

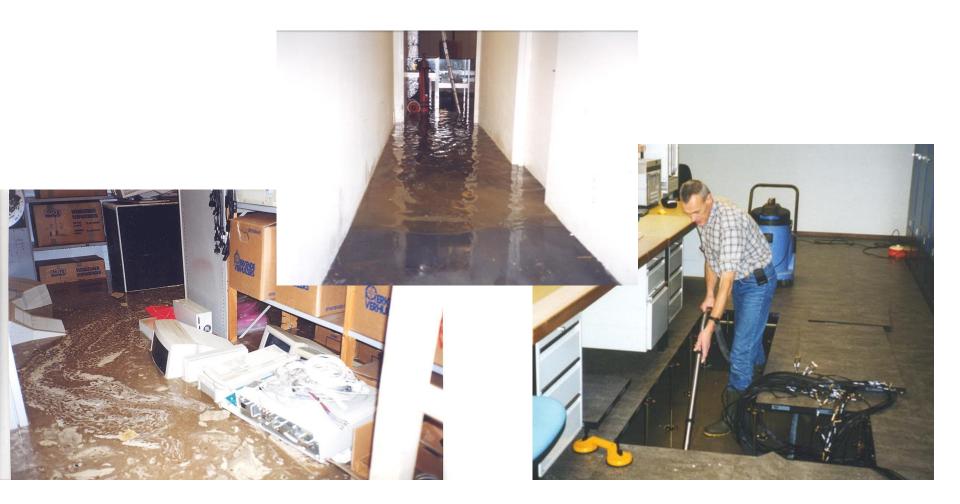
uan Langeveilde, Philistraem, Canusay, Jaffe, Schilitzsi, 2000, A&A 354, L45





How lucky we were

The Biblical Deluge of 28 Oct 1998







Acknowledgements

- Richard Porcas
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